

Forming Ridge-and-Trough Systems on Icy Satellites: Insights from Physical Analogue Modeling

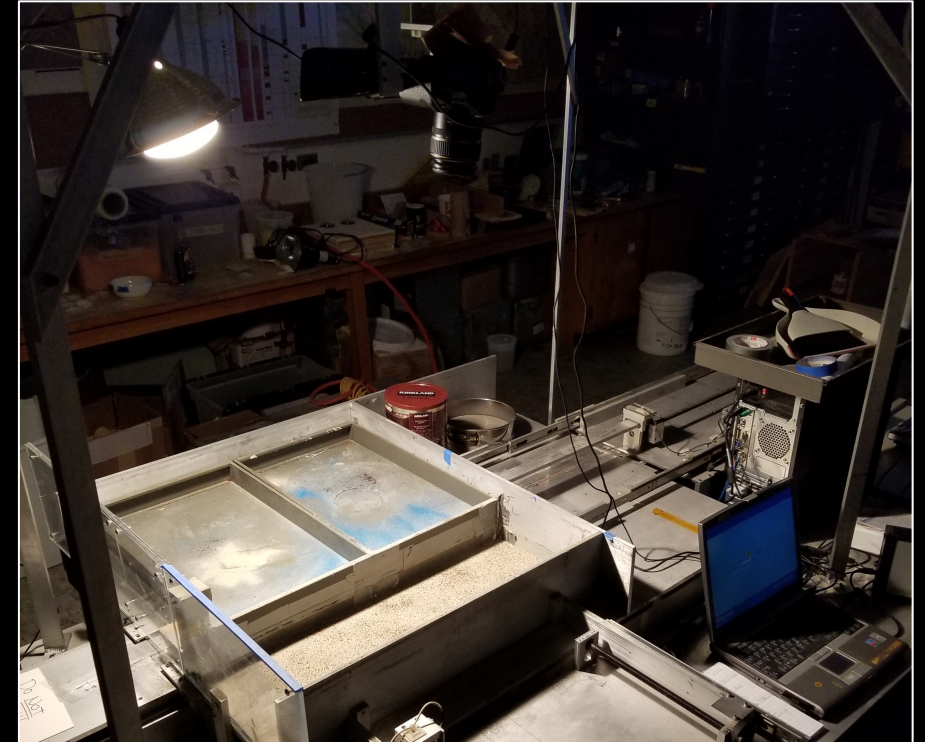
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Overview

- Observations and Motivation
- Analogue Set-up
- Experiment Results
 - No Ductile Layer
 - Brittle layer thickness – compression
 - Ductile layer thickness
 - Extension
 - Scaling to ice shells
- Summary and Future Work

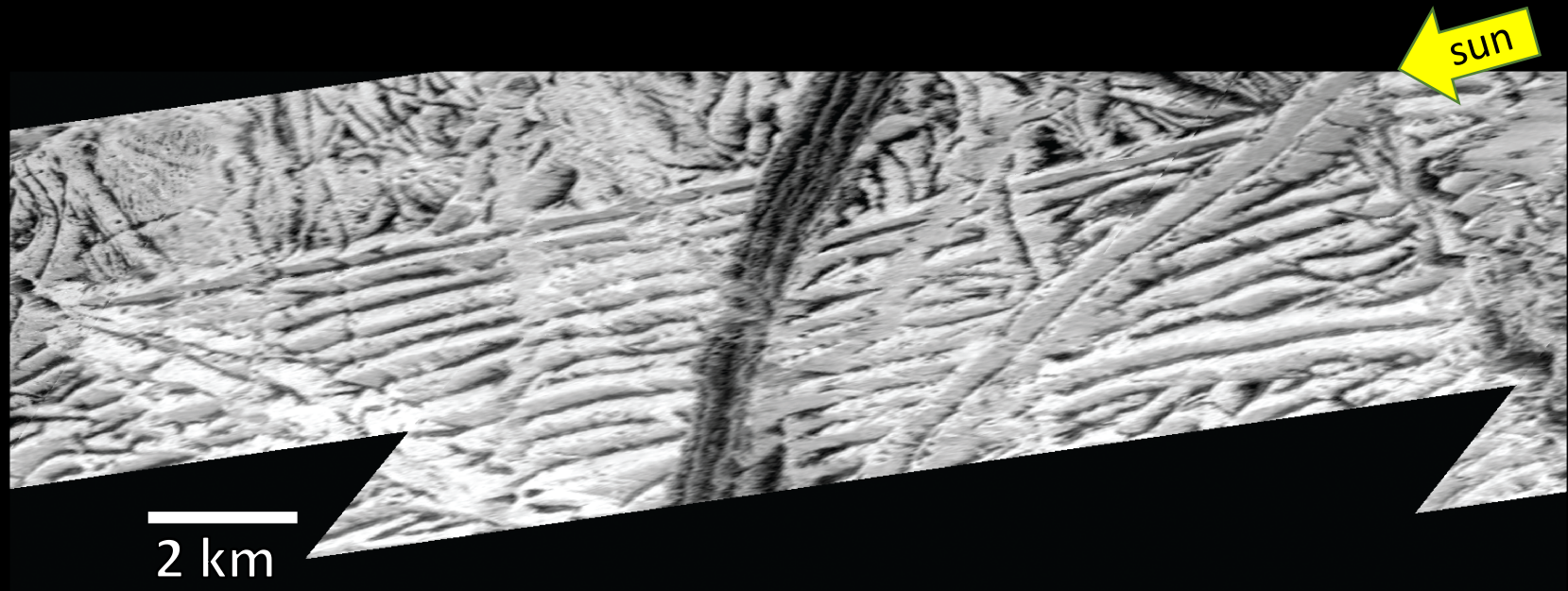
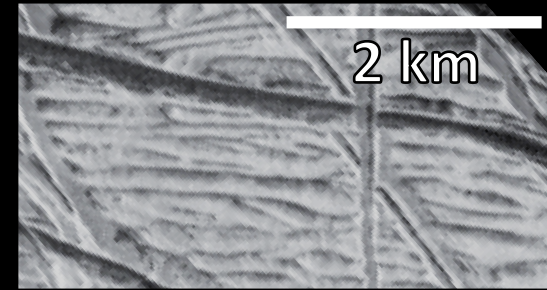
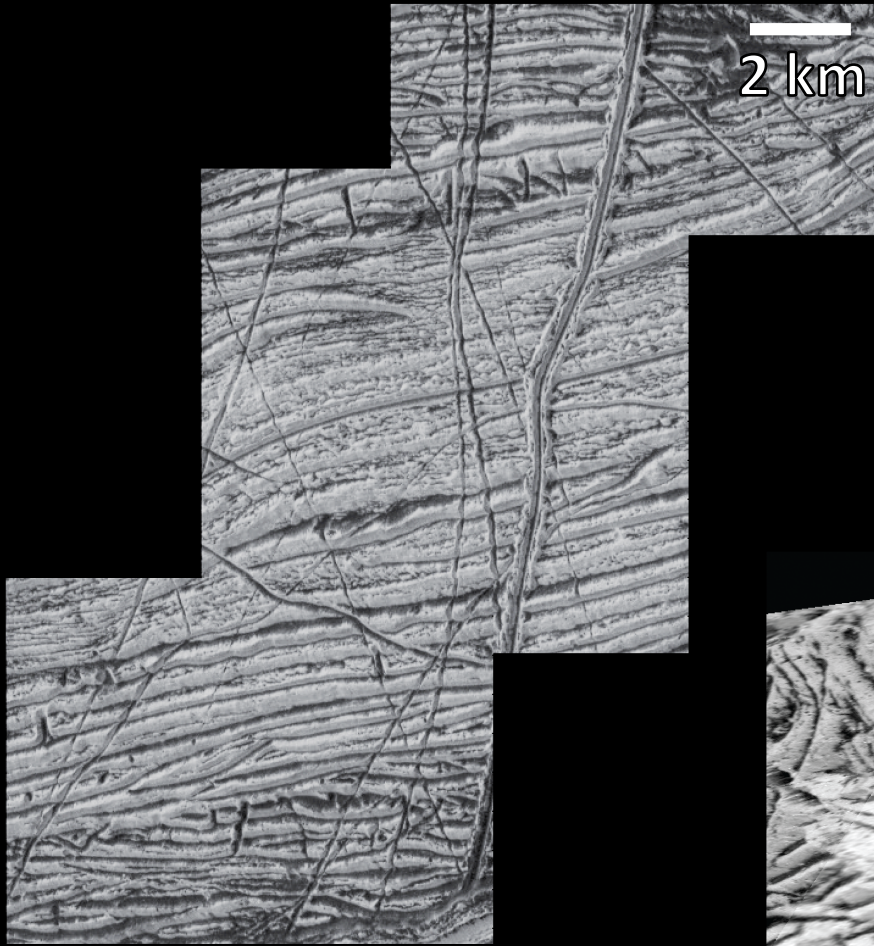


Overview

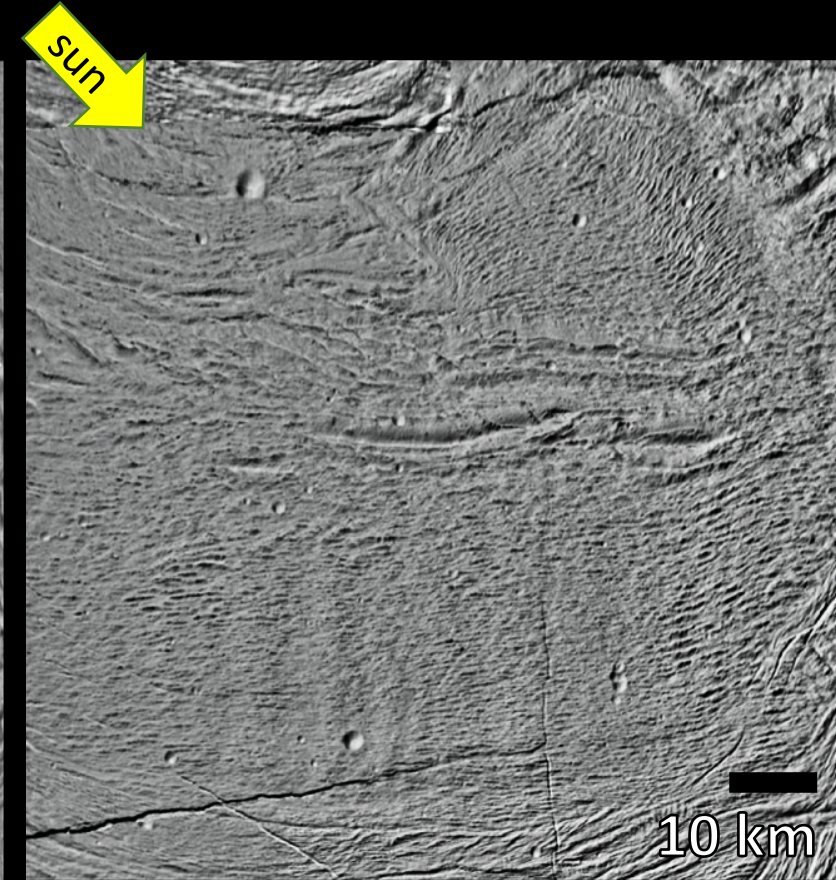
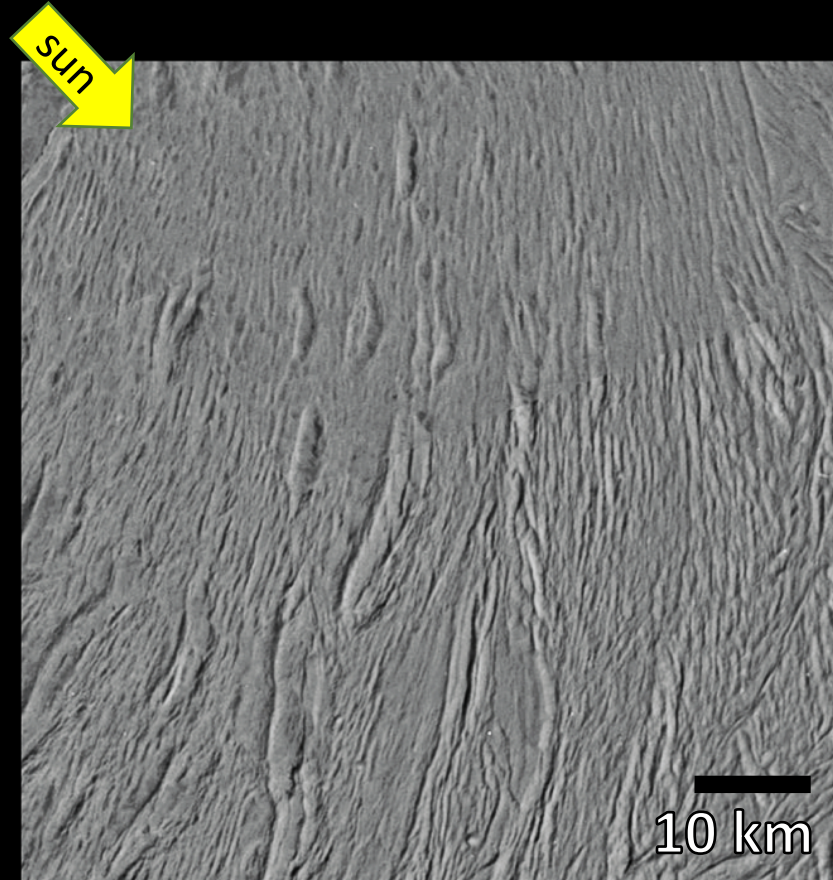
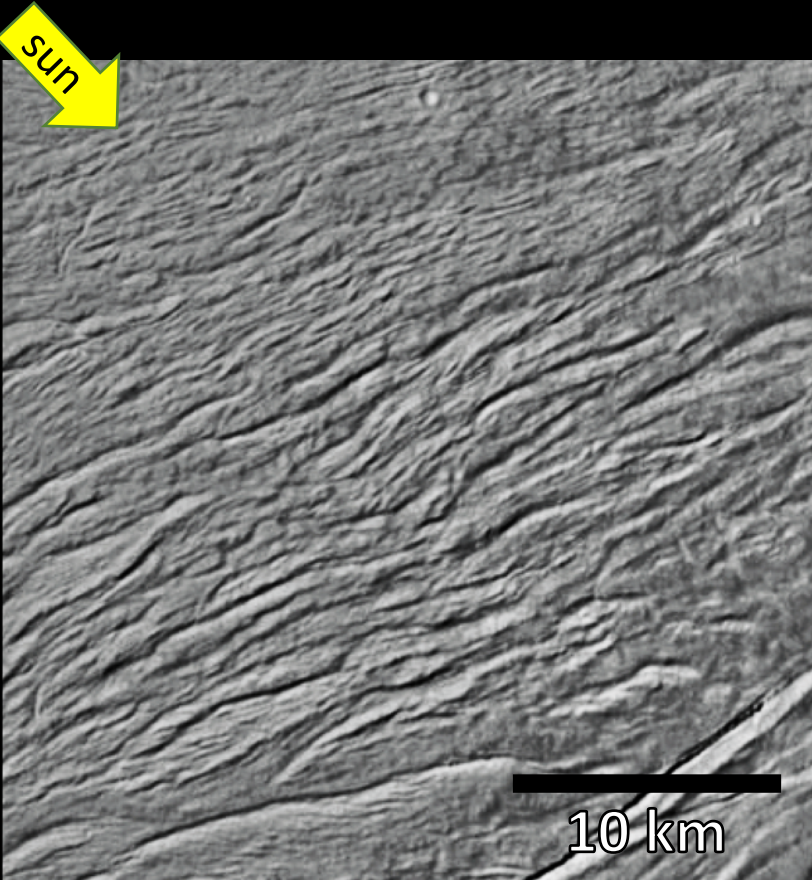
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The ductile subsurface ice and brittle layer thickness are key to the formation of ridge-and-trough systems on icy satellites.

Ridges on Europa and Enceladus



Ridges on Europa and Enceladus



Ridges on Europa and Enceladus

- Ridge-and-trough systems on Europa and Enceladus dominate the surface
 - Europa at 100 meter scale
 - Enceladus at kilometer scale
- How did these ridges form?
 - Compression (e.g., Leonard et al., 2015)? Extension (e.g., Kattenhorn, 2002)?
- What role, if any, did subsurface ductile ice play in the formation process?

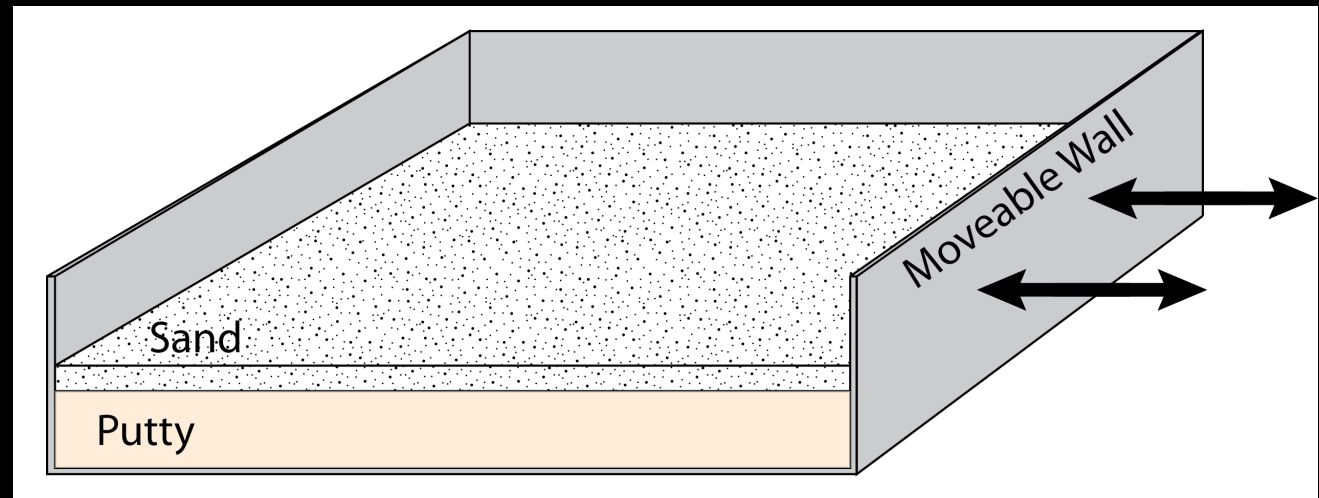
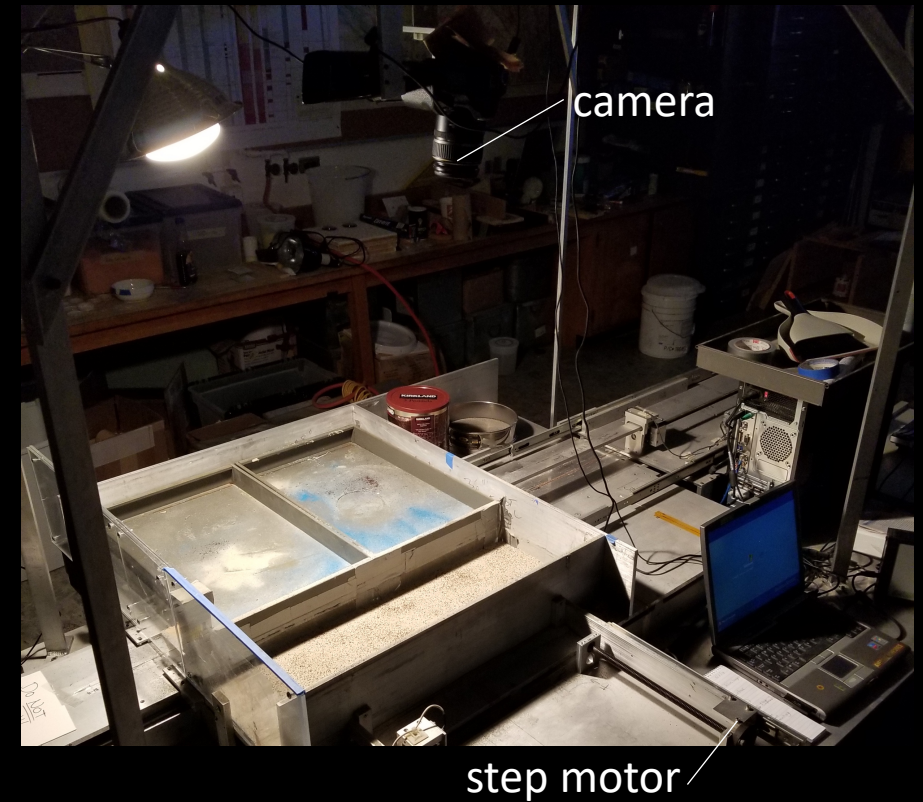
Analogue Experiments

- Brittle layer = sand (~200 μm)
- Scaling:
 - 1 cm model is 2 km on Europa
 - 1 cm model is 20 km on Enceladus

$$\frac{h_{ice}}{h_{mod}} = \frac{\rho_{mod}}{\rho_{ice}} \cdot \frac{g_{mod}}{g_{ice}} \cdot \frac{C_{ice}}{C_{mod}}$$

e.g., Cruz et al. (2007); Hubbert (1937)

- Coffee grounds sprinkled on the surface (passive markers)

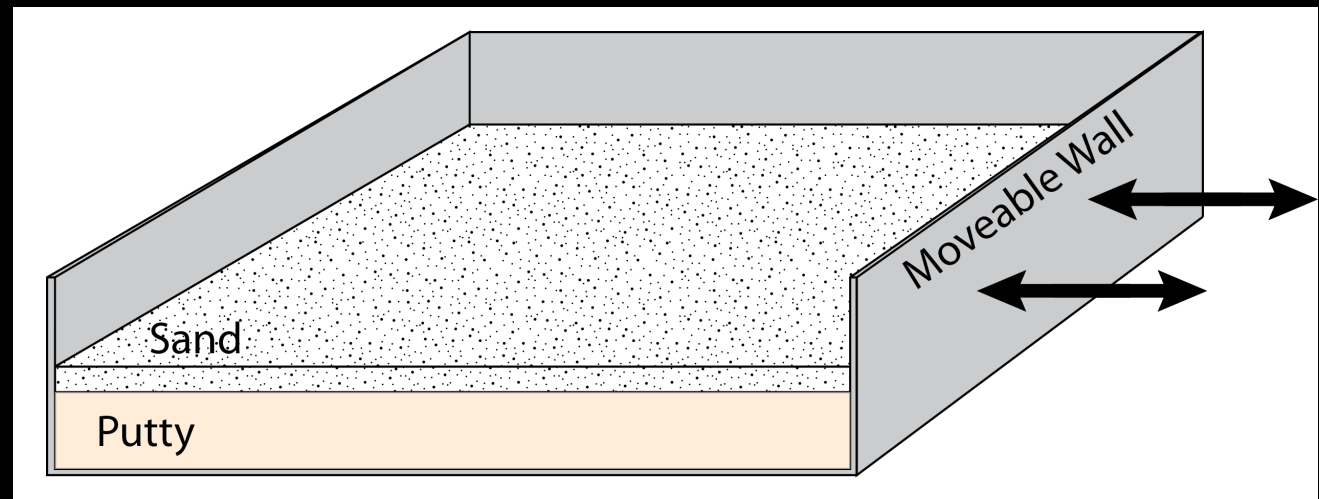
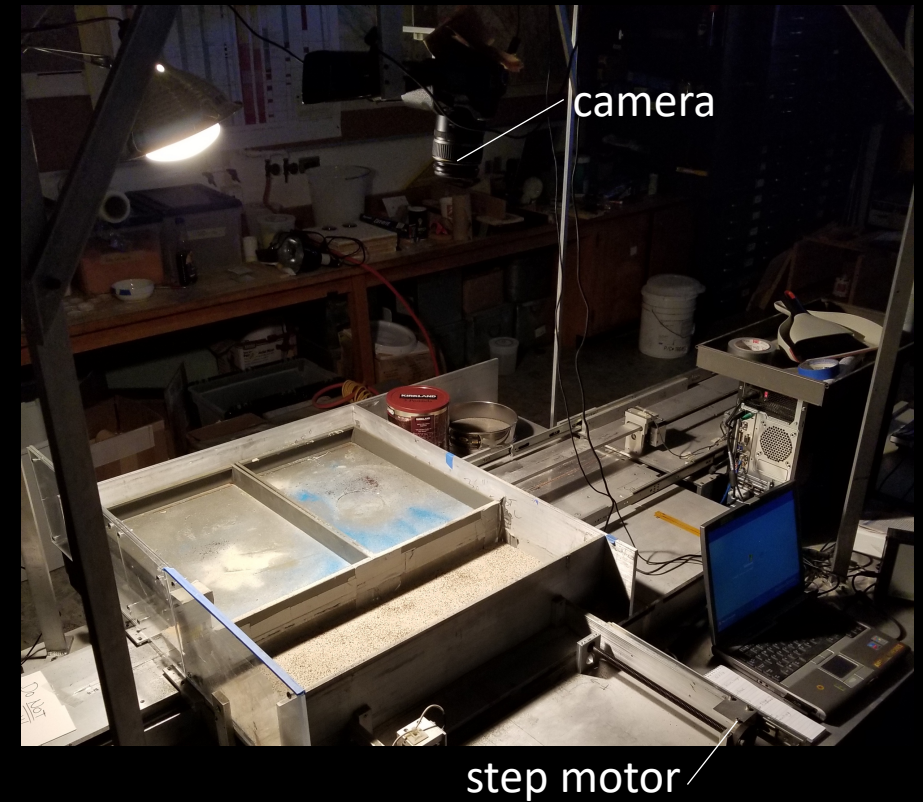


Analogue Experiments

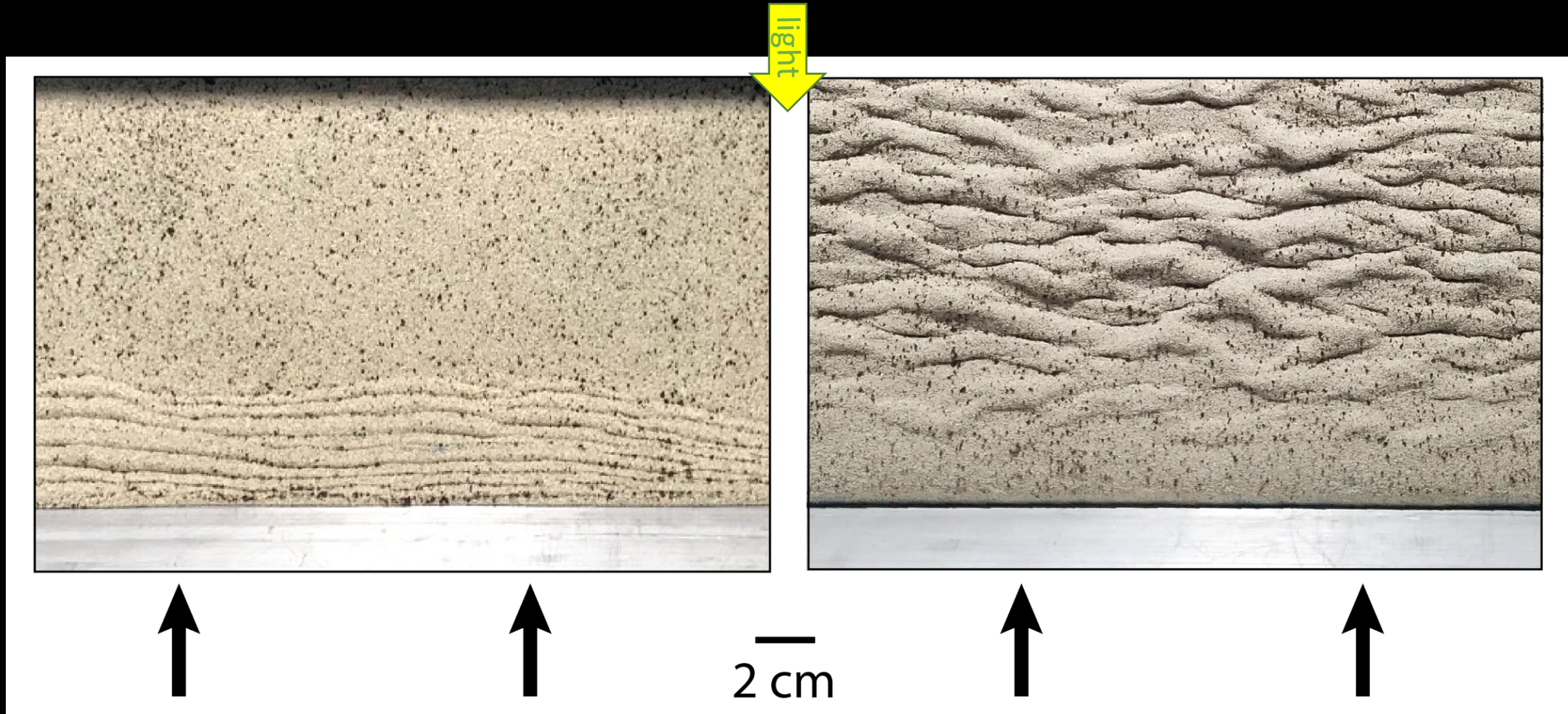
- Ductile layer = putty (10^4 Pa s)
- Scaling:
 - 24 hours model is $\sim 10^6$ years on Europa or Enceladus
 - 10^{-5} s^{-1} model is 10^{-13} s^{-1} on Europa or Enceladus

$$\frac{t_{ice}}{t_{mod}} = \frac{\mu_{ice}}{\mu_{mod}} \cdot \frac{C_{mod}}{C_{ice}}$$

e.g., Davy and Cobbold (1988)



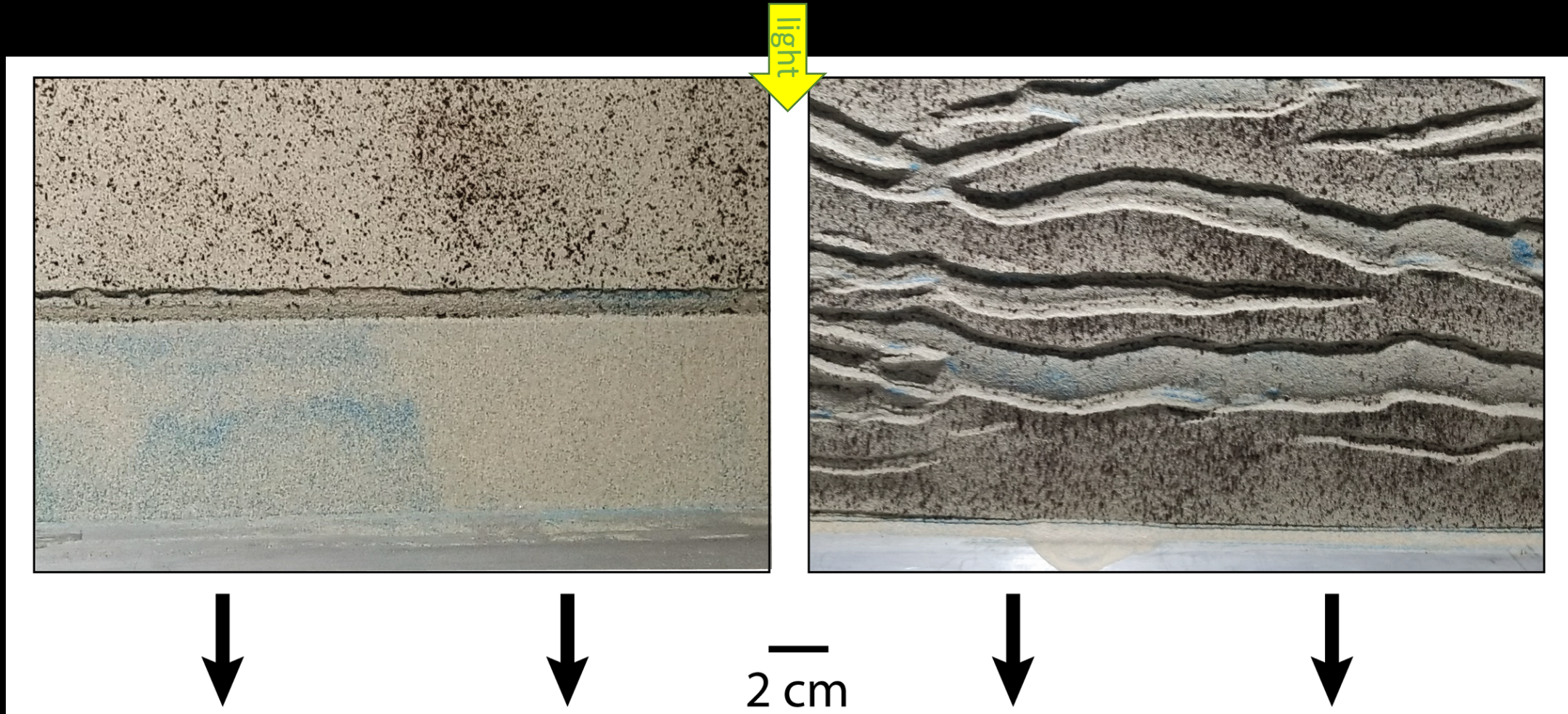
Results – Brittle vs. Brittle+Ductile Tectonics



Sand thickness ~ 0.25 cm

Sand thickness ~ 0.25 cm
Putty thickness ~ 2 cm

Results – Brittle vs. Brittle+Ductile Tectonics

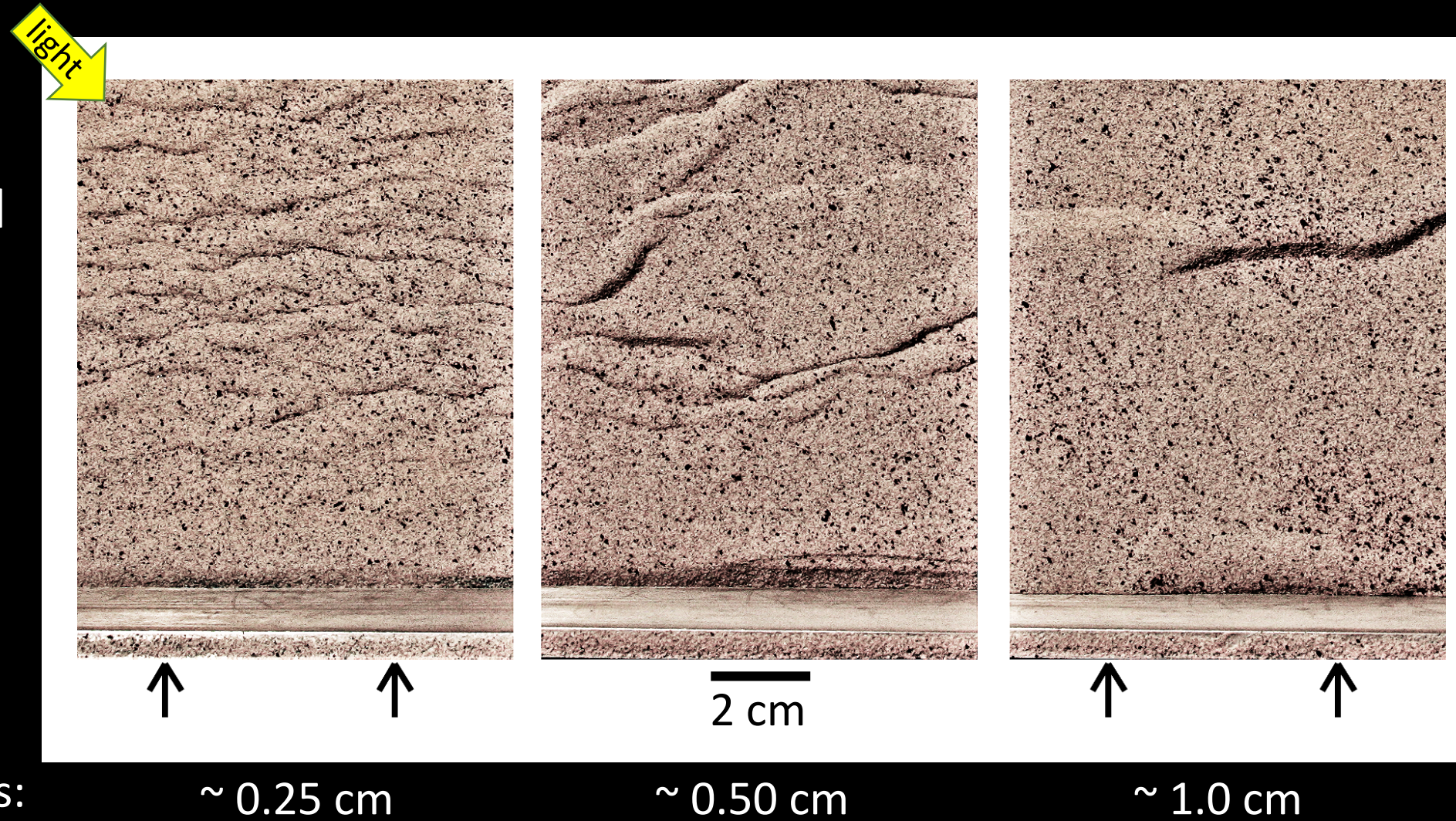


Sand thickness ~1.0 cm

Sand thickness ~1.0 cm
Putty thickness ~2 cm

Results — Compression with Varying Brittle Layer Thickness

- Brittle layer thickness varied
- Putty thickness all ~ 2 cm
- Run for 24 hours, $\sim 33\%$ strain



0.25 cm sand



0.50 cm sand



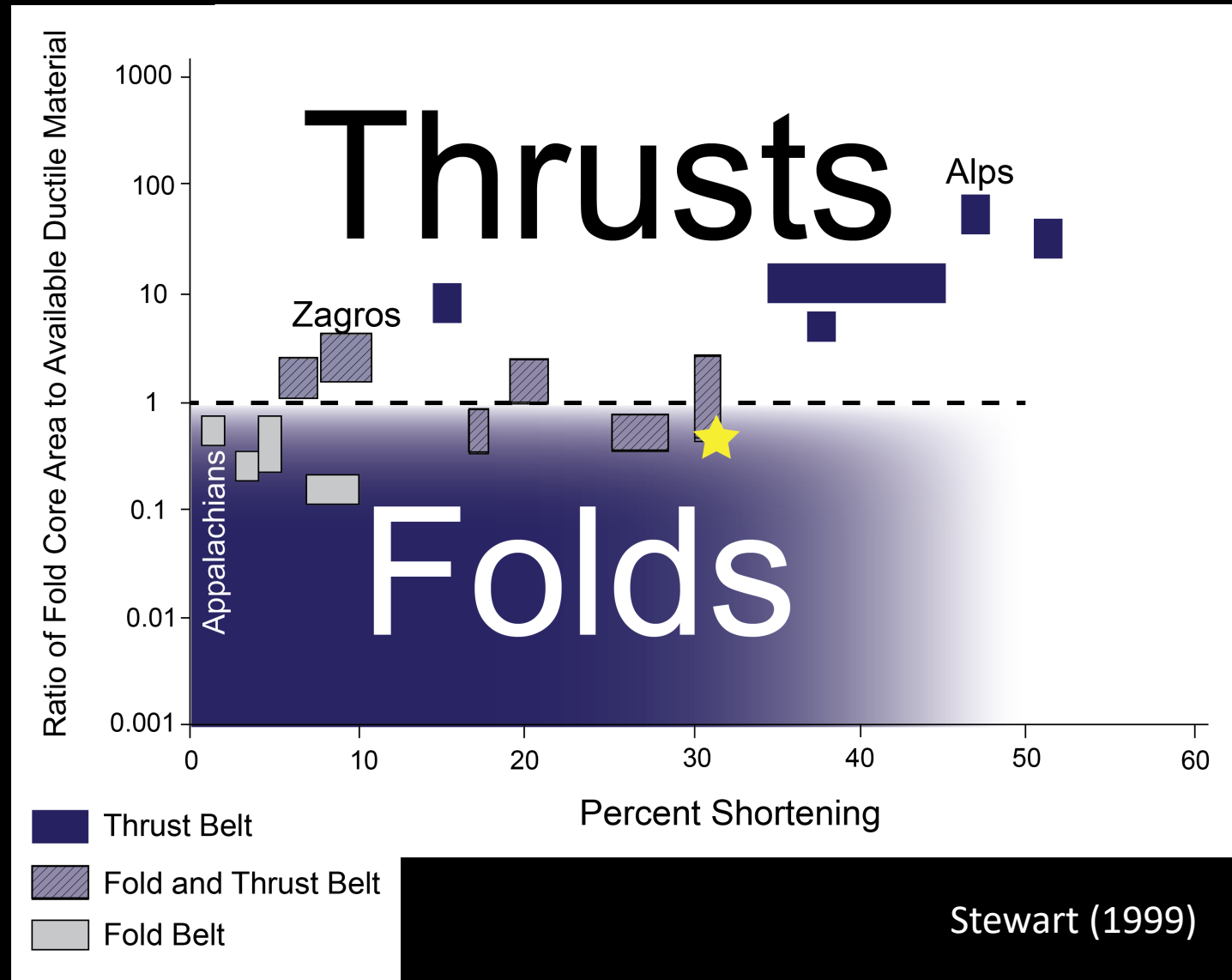
1.0 cm sand




2 cm

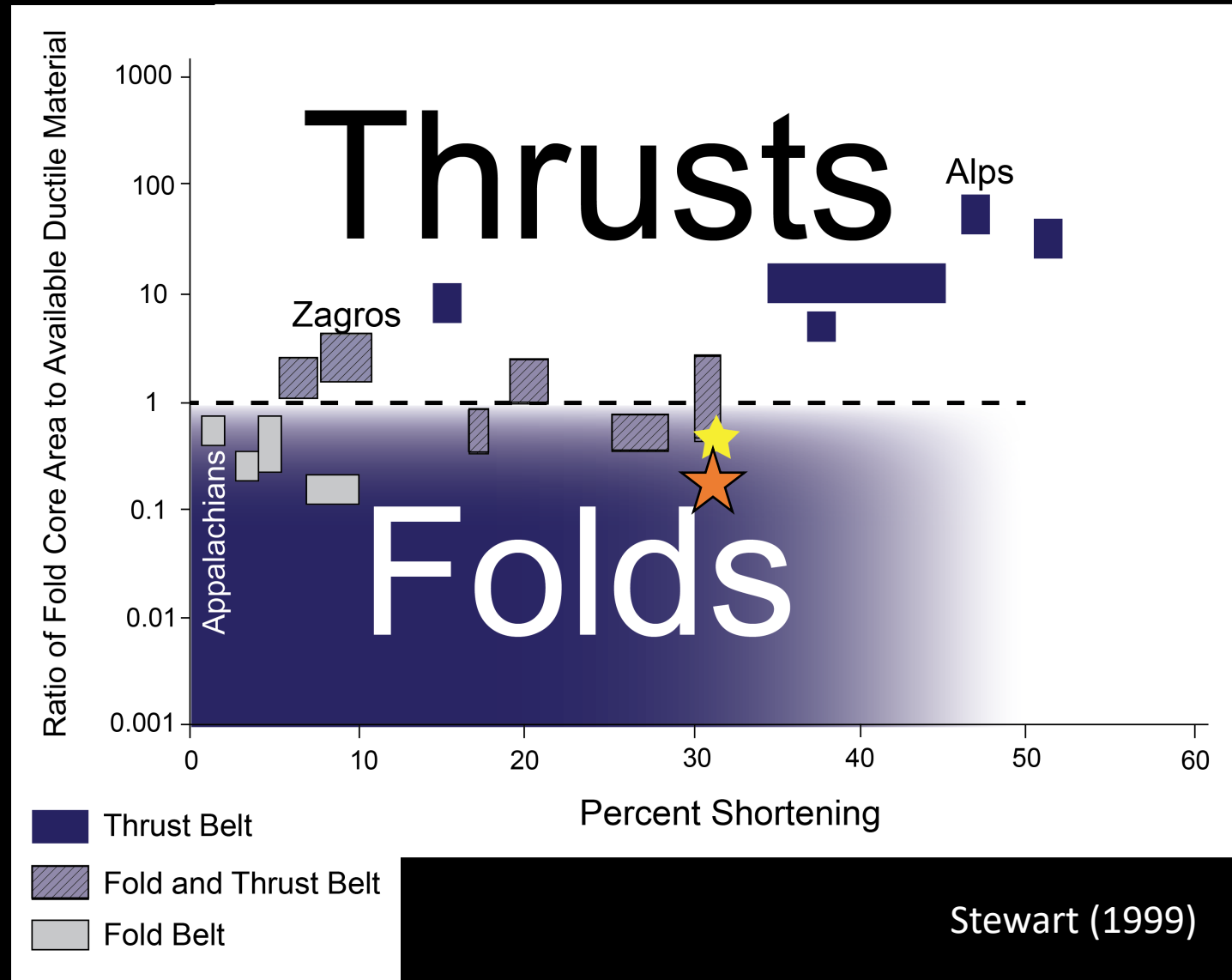
Results – Ductile Layer Thickness (Compression)

- Regime diagram from Stewart (1999)
 - Our compression experiments sit near the border of the transition from thrusts to folds
- The thickness of the ductile layer may be effecting the surface structures



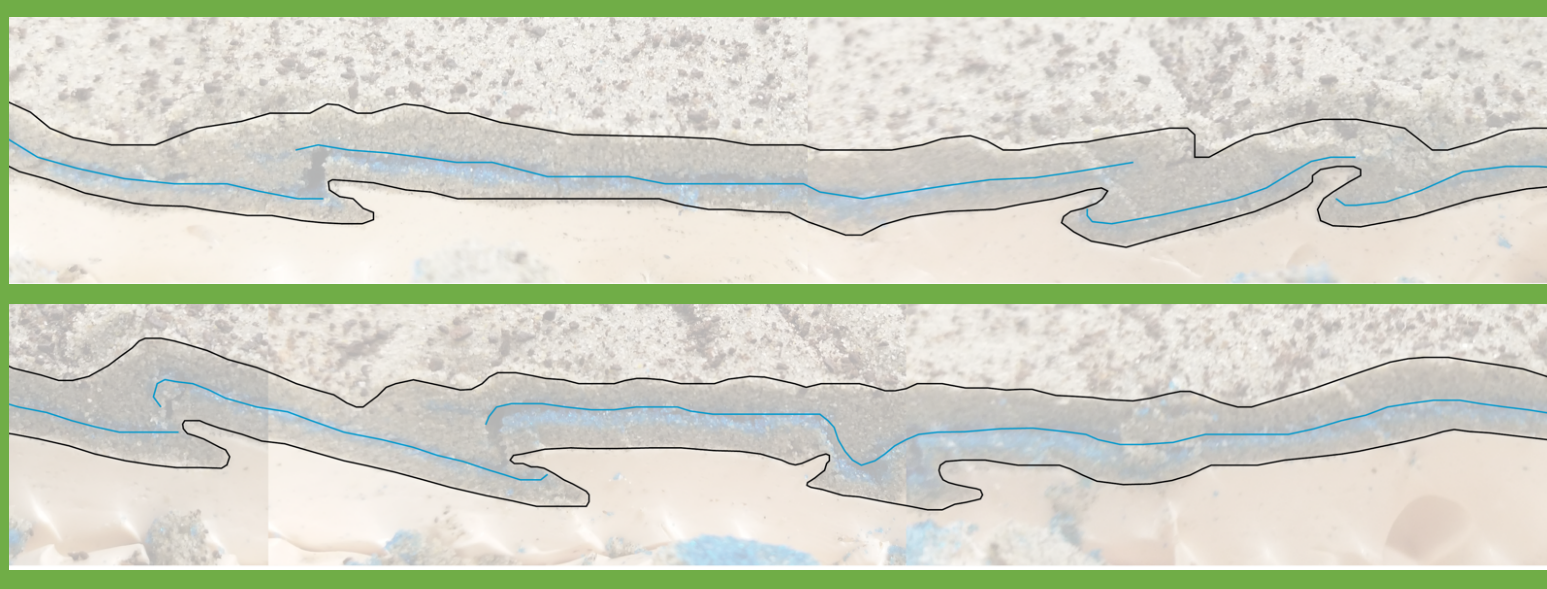
Results – Ductile Layer Thickness (Compression)

- Regime diagram from Stewart (1999)
 - Our compression experiments sit near the border of the transition from thrusts to folds
- The thickness of the ductile layer may be effecting the surface structures
 - Test it!



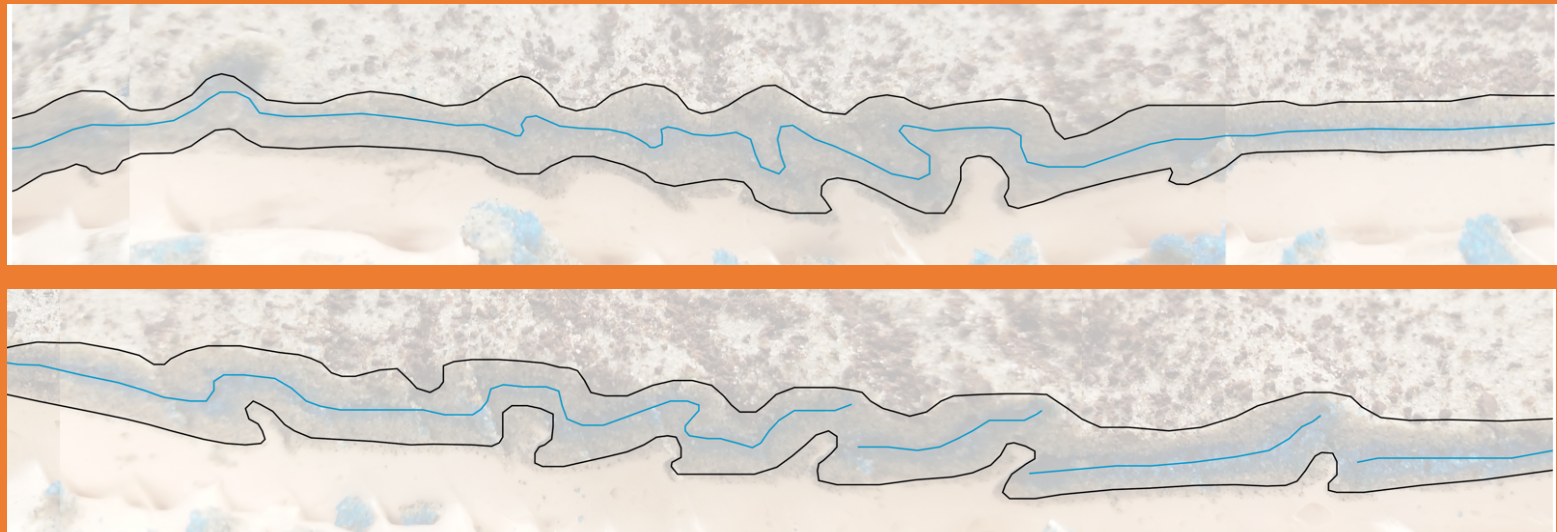
Stewart (1999)

Results – Ductile Layer Thickness (Compression)



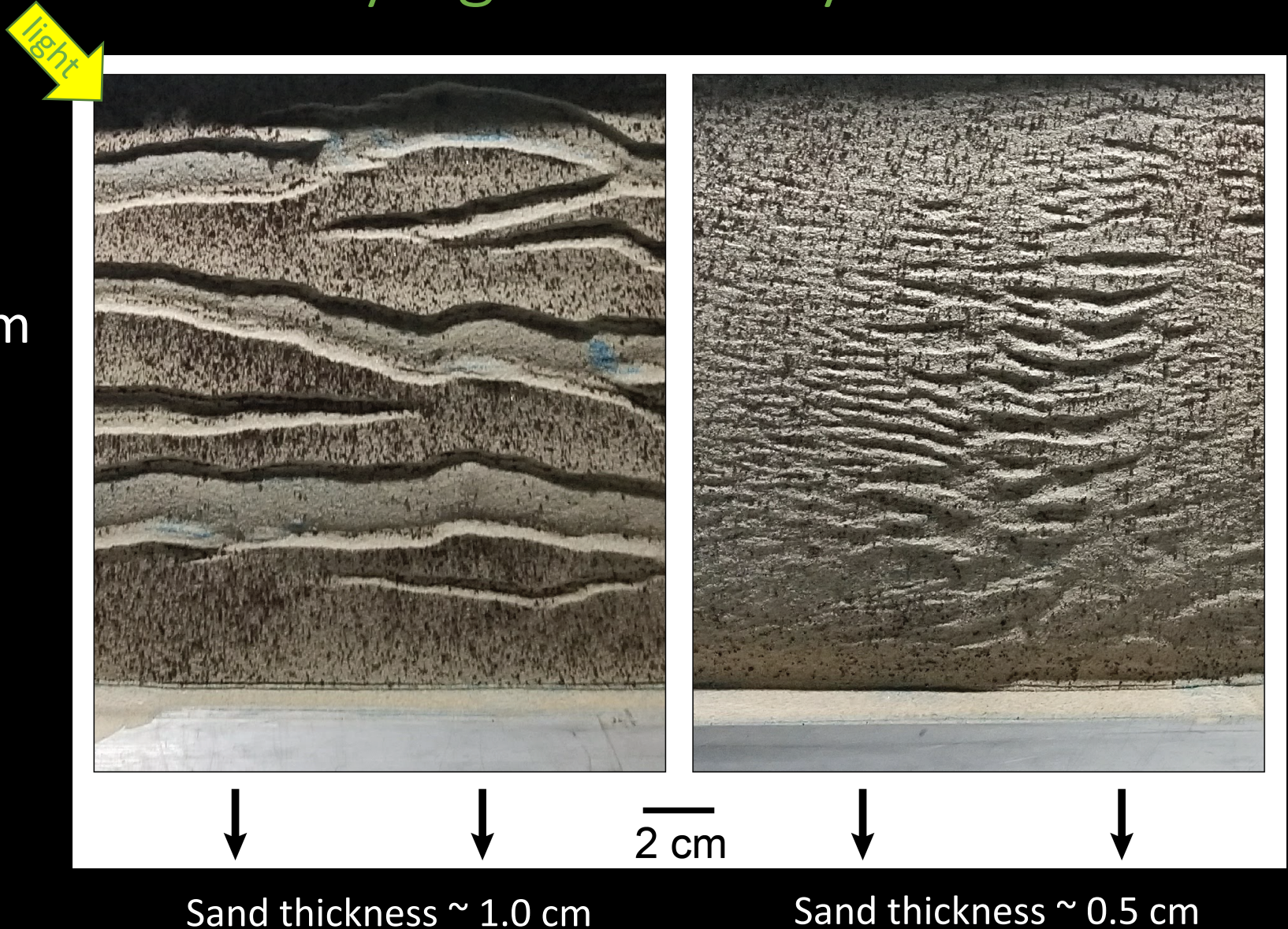
- Putty thickness ~ 2 cm
 - Sand thickness ~ 0.5 cm
- Thrusts (1 - 2 cm wavelength) dominate

- Putty thickness ~ 4 cm
 - Sand thickness ~ 0.5 cm
- Folds (0.5 – 1 cm wavelength) dominate

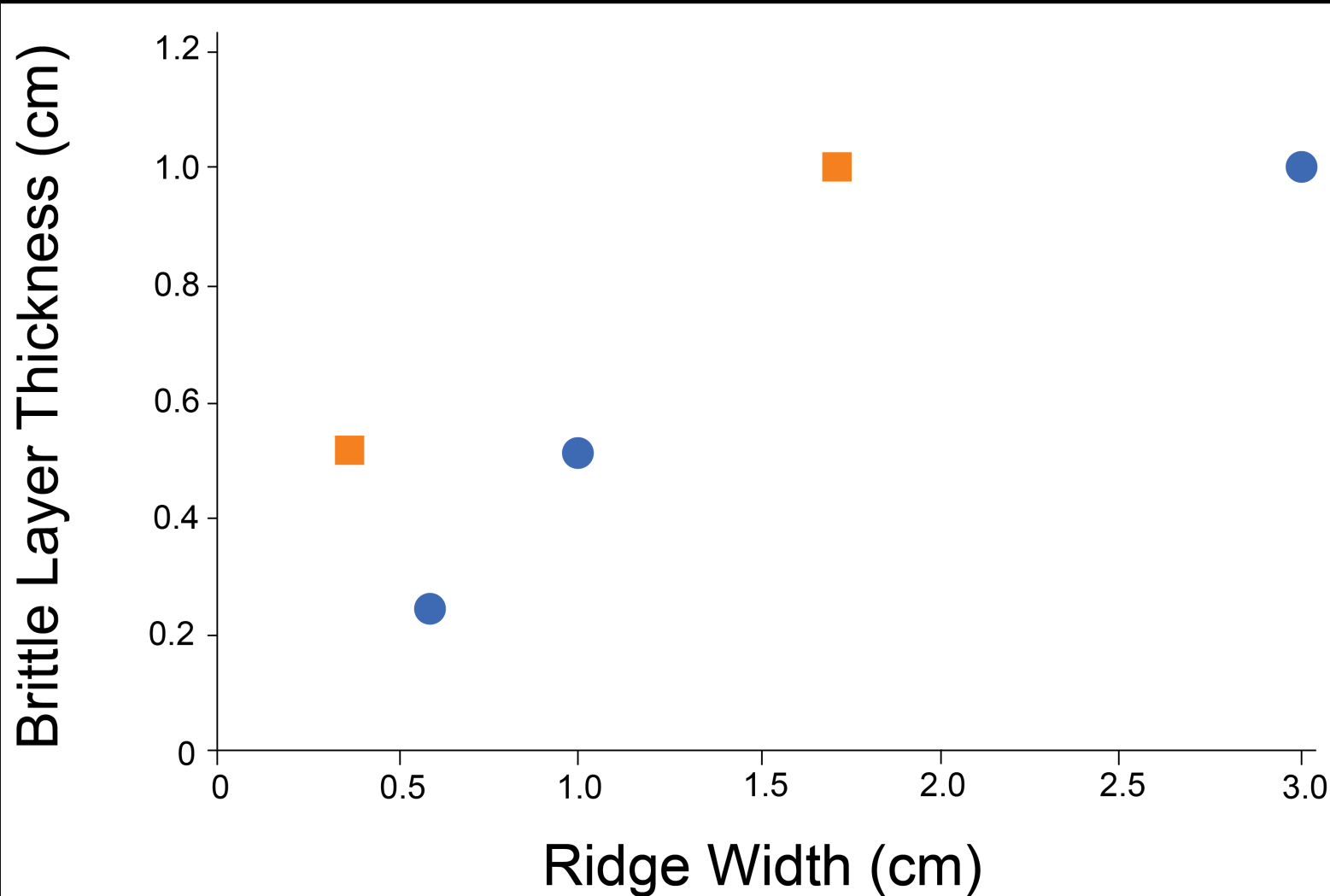


Results – Extension with Varying Brittle Layer Thickness

- Brittle layer thickness varied
- Putty thickness all ~ 2 cm
- Run for 24 hours, $\sim 33\%$ strain
- Tilt-blocks only occur when extension rate is near or at maximum



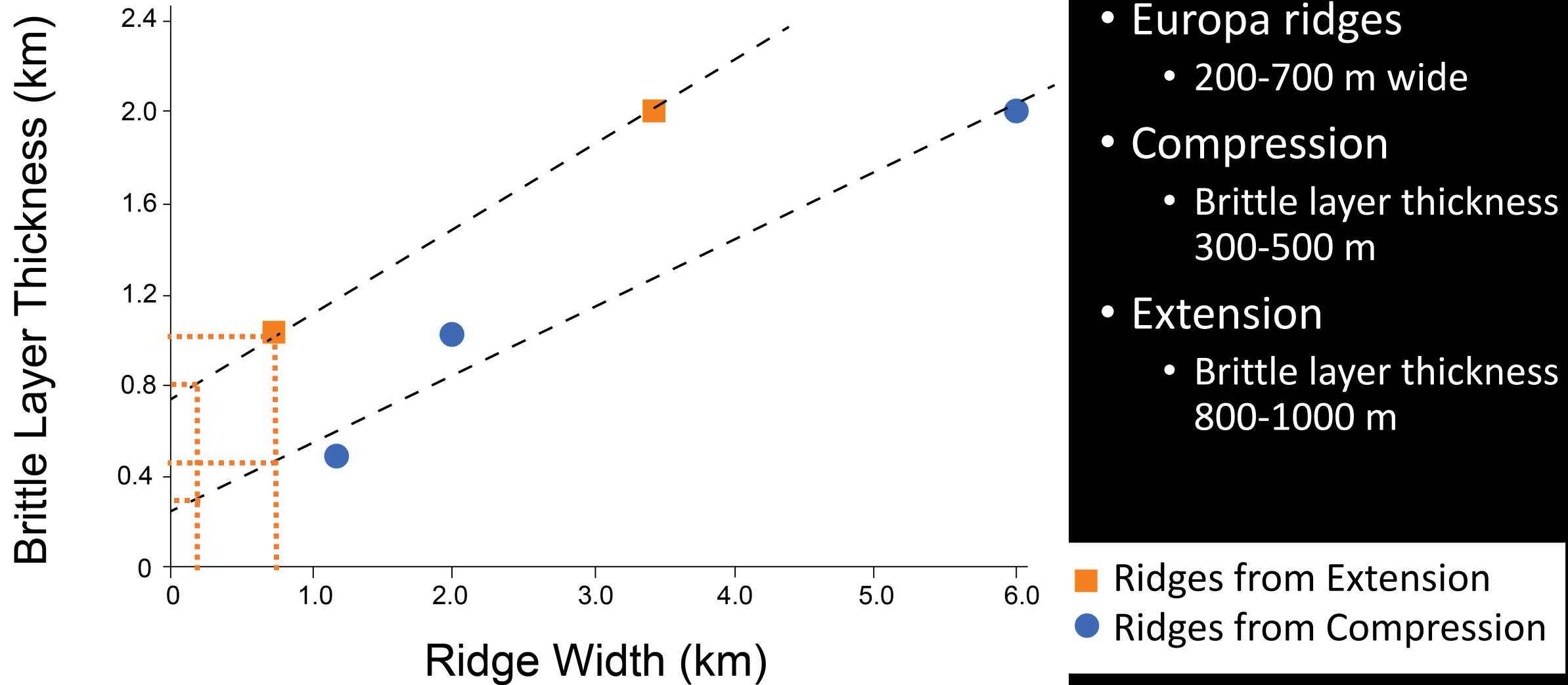
Results – Implications for Icy Satellites



- Ridge width produced by each extension or compression experiment

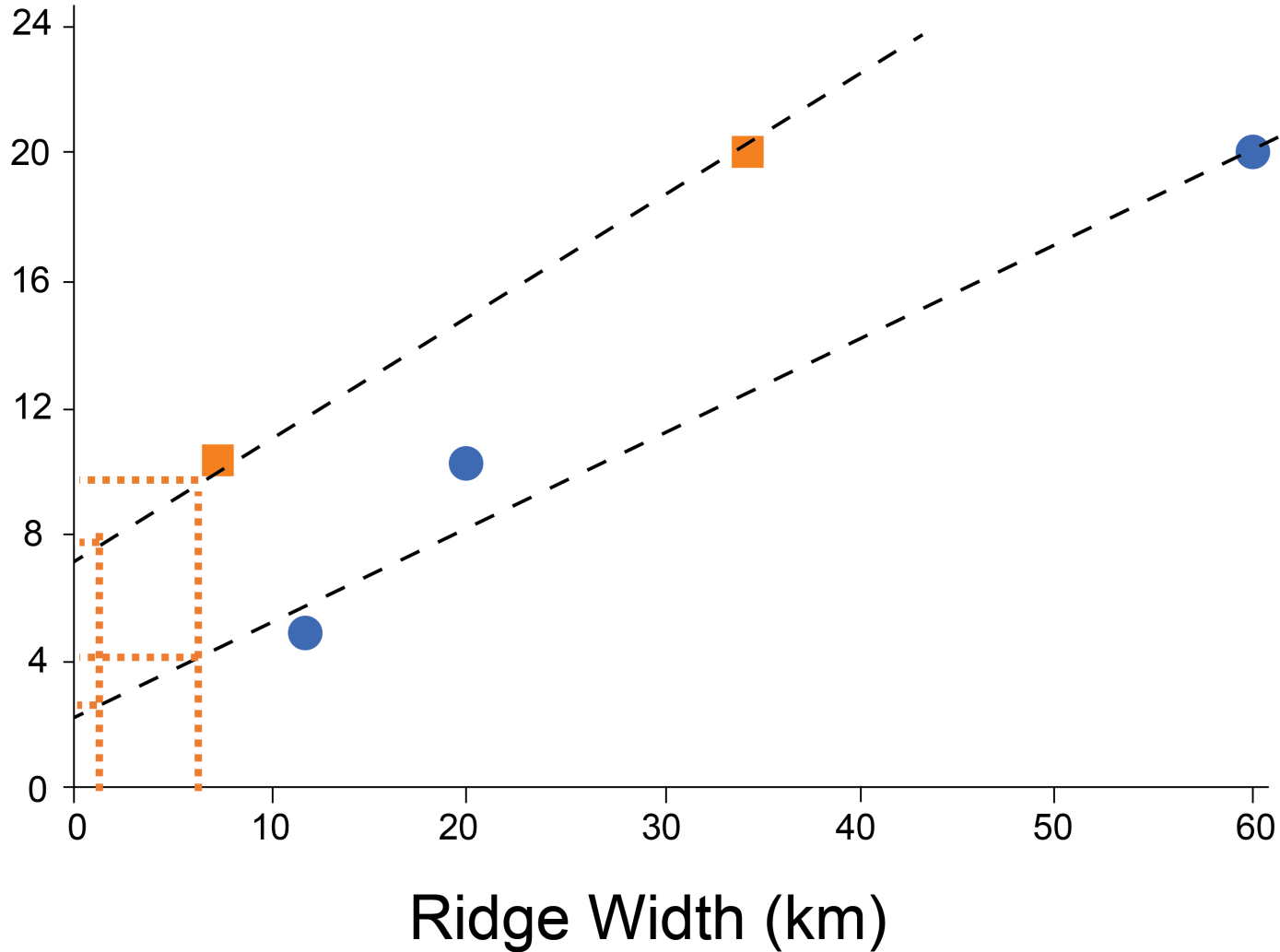
■ Ridges from Extension
● Ridges from Compression

Results – Implications for Europa



Results – Implications for Enceladus

Brittle Layer Thickness (km)

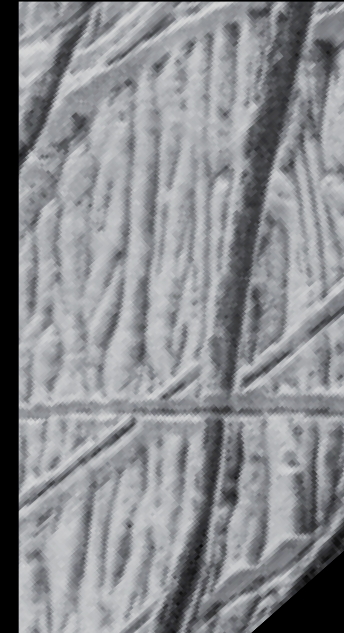


- Enceladus ridges
 - 1-7 km wide
- Compression
 - Brittle layer thickness 2-4 km
- Extension
 - Brittle layer thickness 8-10 km

- Ridges from Extension
- Ridges from Compression

Highlights of Analogue Experiments

- The ductile ice and brittle layer thickness are key to the formation and morphology of ridge-and-trough systems
 - Implications for the evolution of the ice shell
- Total thickness of the ice shell or ductile subsurface ice could determine dominant process (e.g. thrusting vs folding)
- Europa ridge-and-trough terrain implies brittle ice thickness of ~ 0.3 to 1 km; Enceladus, ~ 2 to 10 km
- Further experiments and morphological analysis can constrain this further



2 km



2 cm